What do we Know About the Universe?

Andreas Albrecht UC Davis dept. of Physics Talk at Lowell Observatory October 1, 2016

Work supported by UC Davis and the US Department of Energy

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News



Jack Gunion has been awarded the APS J. J. Sakurai Prize in Theoretical Particle Physics Posted: Sep 28, 2016, 11:05 AM

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New Graduate Fellowship

Posted: Sep 2, 2016, 2:36 PM The James D. Cone Graduate Fellowship has been established through a generous donation...



Charles Fadley was elected as Honorary Member of the The International Science Committee for the International Conferences on Vacuum Ultraviolet and X-ray Physics Posted: Aug 5, 2016, 4:47 PM

The nomination read: "Chuck Fadley is widely regarded as the most inspiring scientist in...



Events

Sep 29, 2016, 3:30 pm Fields, Strings, Gravity - Mark <u>Mezei</u> Sep 29, 2016, 4:10 pm <u>Condensed Matter - Catherine</u> <u>Conlon</u> Sep 30, 2016, 12:15 pm <u>QMAP seminar: Arnab Rudra</u> (<u>MSB 2112</u>) Oct 3, 2016, 1:30 pm Joint Theory Seminar: Gilly Elor Oct 3, 2016, 4:10 pm

Colloquium - David Schwab

Oct 6, 2016, 3:30 pm Fields, Strings, Gravity - Matt von Hippel

Oct 6, 2016, 4:10 pm Condensed Matter - Adrian Swartz

Oct 7, 2016, 12:10 pm Cosmology - Aparna Venkatesan

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UC Davis is ranked 43rd among all U.S. public and private universities in the ranking

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Q

jan shrem and maria manetti shrem museum of art



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A Brand New Museum

Doors Open November 13: Free to All

Grounded in the legacy of UC Davis' world-renowned first generation art faculty, the Jan Shrem and Maria Manetti Shrem Museum of Art will be a hub of creative practice for today's thinkers, makers and innovators, now and for generations to come.

The museum is under construction and is set to open on November 13, 2016. Check back for more information on opening events this fall.

Be a part of making it happen! Here's how:







A. Albrecht @ Lowell 10/1/16

















The Keck 10m Telescopes on Mauna Kea, Hawaii

A. Albrecht @ Lowell 10/1/16





Segments of the Keck 10m Telescope Mirror

Outline

- 1. Introduction (The "Golden age of cosmology")
- 2. The Big Picture
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The APM (Automatic Plate Machine) Survey (1992) Sky positions of 2,000,000 Galaxies

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The Sloan Digital Sky Survey (to locate over 100,000,000 galaxies, 3D positions for 1,000,000)







A simulation of just 65,000 Sloan galaxies

June 5 2001: First release of Sloan data (50,000 galaxies)





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Sloan Survey Status

Imaging (Galaxy positions on the sky)



47% Complete Jun 21 2002 → 47,000,000 galaxy positions

Spectroscopy (3D galaxy positions)



34% Complete Jul 15 2002
→ 340,000 galaxy positions

Sloan Survey Status

Imaging (Galaxy positions on the sky)

97% Complete Jun 27 2004 → 97,000,000 galaxy positions

Spectroscopy (3D galaxy positions)



67% Complete Jun 27 2004 → 670,000 galaxy positions

Sloan Survey Status

Imaging (Galaxy positions on the sky)

107% Complete Mar 13 2005 → 107,000,000 galaxy positions

Spectroscopy (3D galaxy positions)



68% Complete Mar 15 2005
→ 680,000 galaxy positions



Plot of a slice of SDSS galaxies



http://sdss.org

Maps of the microwave sky (the "edge of the observable

universe")









<u>Maps of the microwave sky (the "edge of the observable</u>

universe")



Updated after WMAP announcem ent, Feb 2003





WMAP 3-yr map



WMAP 5-yr map



Maps of the microwave sky (the "edge of the observable

<u>universe")</u>



Updated after Planck announcem ent, 2013



2013



Maps of the microwave sky (the "edge of the observable universe")

Real

1993

Real Dat

COBRAS/SAMB

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D13

<u>Maps of the microwave sky (the "edge of the observable</u> <u>universe"</u>)

Real

1993

Real Dat

March 17 2014! BICEP2 reports

> May 2 2015 Planck reports better polarization data most likely due to nearby dust

> > Real Data!



Maps of the microwave sky (the "edge of the observable

<u>universe")</u>

Real

1993

Real Data!



September 14 2015! LIGO reports direct detection of gravitational waves from two merging black holes





Links related to previous slides

http://www.esa.int/esaSC/120398_index_0_m.html

http://www.rssd.esa.int/index.php?project=planck

http://bicepkeck.org/

http://www.esa.int/spaceinimages/Images/2015/02/Polaris ation_of_the_Cosmic_Microwave_Background

http://www.esa.int/esaSC/120398_index_0_m.html

http://www.rssd.esa.int/index.php?project=planck

http://albrecht.ucdavis.edu/special-topics/bicep2-story

https://www.ligo.caltech.edu/news








Mass inferred from lensing: Must have dark matter







Using Hubble's "advanced camera for surveys" installed June 2002



Galaxy Cluster Abell 1689 Hubble Space Telescope • Advanced Camera for Surveys

NASA, N. Benitez (JHU), T. Broadhurst (The Hebrew University), H. Ford (JHU), M. Clampin(STScI), G. Hartig (STScI), G. Illingworth (UCO/Lick Observatory), the ACS Science Team and ESA STScI-PRC03-01a A. Albrecht @ Lowell 10/1/16 http://hubblesite.org/

http://www.nasa.gov/mission_pages/hubble/main/index.html

Some Future Plans



LSST (Large-aperture Synoptic Survey Telescope)



WFIRST





LSST (Largeaperture Synoptic Survey Telescope)

Some Euro New facilities being built

LSST (Largeaperture Synoptic Survey Telescope)







LSST (Large-aperture Synoptic Survey Telescope)



James Webb Space Telescope (2018 Launch)



Frequently Asked Questions

1. Will the WFIRST mission be a breakthrough in the search for dark matter?

WFIRST will survey large areas of the sky measuring the effects of dark matter on the distribution of galaxies in the universe. It will also observe distant Type Ia supernovae to use them as tracers of dark matter and dark energy. It will provide a huge step forward in our understanding of dark matter and dark energy.

2. In what phase of development is currently the WFIRST spacecraft?

WFIRST is currently in Phase A.The purpose of Phase A is to develop the mission requirements and architecture necessary to meet the programmatic requirements and constraints on the Project and to develop the plans for the Preliminary Design phase.

3. Are the preparations on track for the mid-2020 launch?

preparations are on track for a mid-2020 launch. Yes, th

WFIRST

<u>https://www.lsst.org/</u>

http://jwst.nasa.gov/index.html

http://wfirst.gsfc.nasa.gov/

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Distances in the Universe

Measure of distance: One Kilometer ≈ Walk from Lowell Observatory to Thorpe Park Tennis courts



Measure of distance: One Kilometer ≈ Walk from Lowell Observatory to Thorpe Park Tennis courts



Count cosmic distances as grains of sand: One grain of sand per kilometer.

Grain of sand (enlarged)



Diameter of earth = 12,760 kilometers $\leftarrow \rightarrow$ 1 Teaspoon of sand





Distance to Moon = 356,410 kilometers ←→ 1 Handful of sand







Distance to Moon = 356,410 kilometers ←→ 1 Handful of sand

(Also roughly the distance light travels in one second)







Distance from Earth to Sun = 149,600,000 kilometers (8 light minutes) $\leftarrow \rightarrow$ 1 Milkshake cup of sand







Distance from Earth to Pluto = 6,000,000,000kilometers $\leftarrow \rightarrow 1$ wheelbarrow of sand





Distance from Earth to Nearest Star = 40,000,000,000 kilometers $\leftarrow \rightarrow 1$ dumpster of sand



Distance from Earth to Edge of our galaxy = 1,000,000,000,000,000,000 kilometers $\leftarrow \rightarrow 1$ Physics/Geology Bulidng full of sand







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Average distance between galaxies = 3×10^{19} kilometers $\leftarrow \rightarrow 1$ baseball stadium full of sand







Farthest visible "object" in the universe: 1×10^{23} kilometers $\leftarrow \rightarrow$ mountain range of sand







What we know about the big picture

1) On large scales the matter in the Universe is spread out very smoothly ("Homogeneous") Mean density: 10^{-29} gram / cm^3 2) The Universe is expanding Distance v = Hr4Hubble law: 3*m* / sec 100*lightyears* Expansion H =Speed A. Albrecht @ Lowell 10/1/16 66

The homogeneity of the Universe



Isotropy of the microwave background (from the "edge of the observable universe") to one part in 100,000





The Hubble law



$$v = Hr$$

$$H = \begin{pmatrix} 3m / \sec \\ 100 light years \end{pmatrix}$$
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Α.

Hubble Expansion



Hot, Dense past

osmology and High Physics nergy Today C ME 000 Expa Atomic . ~10eV Physics Nuclear nSIO ~ 10 N/e V ng ~/GeV Weak Interactions Tel n man Unificatio Super strings Lowell 10/1/16 Α. 0 M
The History of the Universe

Time



The History of the Universe

Time



The History of the Universe

Time



Acceleration of the universe



The Hubble law at great distances' depends on the variations of the Hubble "constant" H with time.

Using supernovae (exploding stars) as cosmic "mileposts", acceleration of the Universe has been detected.







Cosmic acceleration (newest data)

Using supernovae (exploding stars) as cosmic "mileposts", acceleration of the Universe has been detected.



Accelerating matter is required to fit current data



Accelerating matter is required to fit current data



Accelerating matter is required to fit current data



Dark Energy and the fate of the Universe



In the presence of dark energy, the simple connection between open/closed/flat and the future of the universe no longer holds









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Cosmic Inflation

- A period of accelerated expansion in the very early universe
- Motivated by particle physics (related to the recently discovered Higgs particle).
- In most models inflation operates when the temperature was 10²⁵ times greater than today!
- Conceptually similar in some ways to the acceleration observed today (interesting relationship between the two)



- Cosmic inflation creates features in the universe on all these different lengths.
- The yellow boxes give the time between "feature creation" in units of 10⁻⁴⁰ seconds!



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Cosmic Microwave Background (CMB) map produced by the Planck satellite (sphere shown using a projection, like in an atlas)

The map shows minute variations in the temperature (just 1 part in 100,000, or in the 5th decimal place).



This plot shows one way to quantify the feature in the CMB map. Roughly, the x-axis labels patch size, and the y-axis show how strongly the temperature typically varies among patches of that size.

Using the CMB to learn about the Universe



- solid=inflation model
- dashed=defect models
- (magenta=desperate)

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- Extraordinarily successful predictions of features in the observed universe
- Very problematic aspects emerge when we attempt to complete the picture. (The cause of intensive research and debate among the experts.)

 May cosmologists believe in "eternal inflation" (our universe exists in a "pocket" with eternal inflation all around us).



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- Eternal inflation theory predicts infinitely many pocket universes, some like ours, some different



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This question appears to lead to deep ambiguities and problems with the theory that cause some to reject the idea of cosmic inflation altogether But Which one is really ours?

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Extraordinarily successful predictions of observed universe

A very exciting place to be!

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Cosmic Inflation



Extraordinarily **successful** predictions of observed universe

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- The cosmic acceleration observed today has proven very difficult to incorporate into our fundamental theories of physics.
- These difficulties have caused some theorists to embrace the "string theory landscape"
- Instead of the physical world around us exhibiting "the fundamental laws", according to the STL picture the universe is made of a landscape of different "worlds" which with their own laws of physics.



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Conclusions

- The search for a "big picture" of the Universe that explains why the region we observe should take this form has proven challenging, but has generated exciting ideas.
- We know we can do science with the Universe
- It appears that there is something right about cosmic inflation
- dSE cosmology offers a finite alternative to the extravagant (and problematic) infinities of eternal inflation
- Predictions of observable levels of cosmic curvature from dSE cosmology will give an important future test

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Conclusion

Amazing data and facilities

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