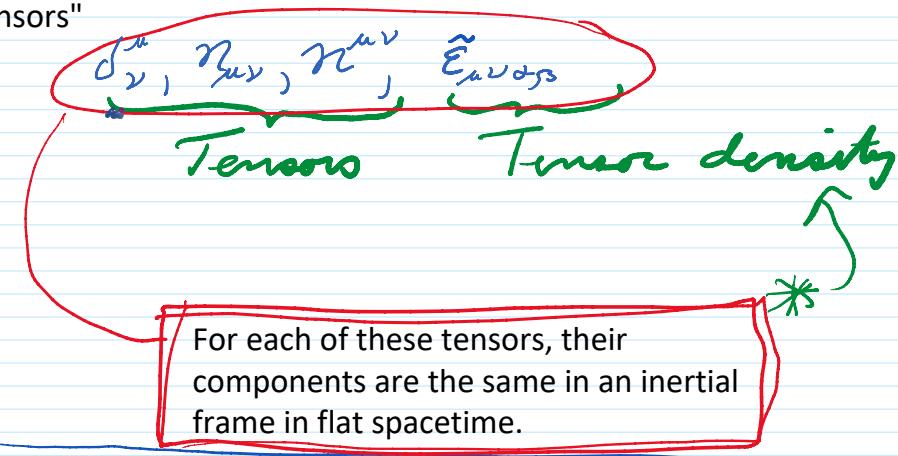


Lecture 4a

Wednesday, January 15, 2020 4:17 PM

The "tensors"



$$\delta^{\mu}_{\nu} = \begin{pmatrix} 1 & & & \\ & 1 & & \\ & & 1 & \\ 0 & & & 1 \end{pmatrix} \quad \text{Identity}$$

used to define inverse

$$A^{\mu\nu} B_{\nu\rho} = \delta^{\mu}_{\rho}$$

$$\Rightarrow B_{\nu\rho} = (A^{\mu\nu})^{-1}$$

$$\begin{aligned} n_{\mu\nu} &\text{ defines invariant } (\delta_{\mu\nu}) \\ (\Delta S)^2 &= n_{\mu\nu} \Delta x^{\mu} \Delta x^{\nu} \\ &= n_{\mu\nu} \Delta x^{\mu} \Delta x^{\nu} \\ &\text{same } n \end{aligned}$$

Note: Any $(0,2)$ tensor $A_{\mu\nu}$
will give a scalar $A_{\mu\nu} \Delta x^{\mu} \Delta x^{\nu}$

will give a scalar $\eta_{\mu\nu} \omega^\mu \omega^\nu$
 but in general the components
 of A will be different in different
 frames.

Also by convention, $\eta_{\mu\nu}$
 is used to "lower an index"
 Given $A^{\mu\nu}$, $A^\mu_{\nu} \equiv \eta_{\nu\rho} A^{\mu\rho}$

(2,0) (1,1)

Again, any (0,2) tensor
 can be used to convert a
 (0,2) to a (1,1), but
 $\eta_{\mu\nu}$ is special because it
 looks the same in all frames.

(Note: δ^μ_ν is a (1,1) rank
 tensor, so it is not a contender)

$\eta_{\mu\nu}$ also defines the "dot" or "scalar"
product:

$$V \cdot W = \eta_{\mu\nu} V^\mu W^\nu$$

and thus the norm

Norm of V :

$$\sqrt{V^\mu V^\nu} \begin{cases} < 0 & V^\mu \text{ is timelike} \\ = 0 & \text{" lightlike (null)} \end{cases}$$

$\eta_{\mu\nu} \sqrt{v^\mu v^\nu}$ {
 $\leq 0 \vee \rightarrow$ universe
 $= 0 \quad " \quad "$ lightlike (null)
 $> 0 \quad v^\mu \text{ is spacelike}$

Scalar \Rightarrow invariant under Lorentz Transf.: -

Note: Norm of $X = (1, 1, 0, 0) = 0$!

$\eta^{\mu\nu}$

The inverse of $\eta_{\mu\nu}$:

(2,0)

$$\eta_{\mu\nu} \eta^{\nu\rho} = \delta_\mu^\rho$$

Used for raising an index

(only (2,0) the same in all frames.)

$\tilde{\epsilon}_{\mu\nu\rho\sigma}$

"Levi-Civita Symbol"

The "completely" antisymmetric with $\tilde{\epsilon}_{12\dots n} = 1$

In 2d: $\tilde{\epsilon}_{ij} = \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix}$ $\tilde{\epsilon}_{12}$

In 3d harder to write out,

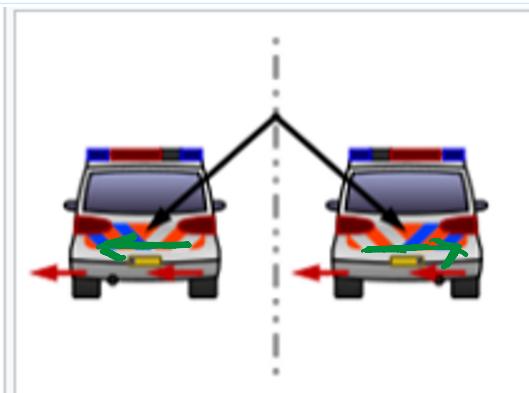
In 3d harder to write out,
but we know it from the
cross product:

$$\begin{aligned}\mathbf{a} \times \mathbf{b} &= a_1 b_1 \mathbf{0} + a_1 b_2 \mathbf{k} - a_1 b_3 \mathbf{j} \\&\quad - a_2 b_1 \mathbf{k} + a_2 b_2 \mathbf{0} + a_2 b_3 \mathbf{i} \\&\quad + a_3 b_1 \mathbf{j} - a_3 b_2 \mathbf{i} + a_3 b_3 \mathbf{0} \\&= (a_2 b_3 - a_3 b_2) \mathbf{i} + (a_3 b_1 - a_1 b_3) \mathbf{j} + (a_1 b_2 - a_2 b_1) \mathbf{k}\end{aligned}$$

$\hat{\mathbf{e}}_1$ $\hat{\mathbf{e}}_2$ $\hat{\mathbf{e}}_3$

$$= \hat{\mathbf{e}}_i a_j b_k \hat{\epsilon}_{ijk}$$

$\mathbf{a} \times \mathbf{b}$ is a pseudovector



Each wheel of the car on the left driving away from an observer has an angular momentum pseudovector pointing left. The same is true for the mirror image of the car. The fact that the arrows point in the same direction, rather than being mirror images of each other indicates that they are pseudovectors.

$\Rightarrow \tilde{\epsilon}_{ijk}$ is pseudotensor
(if it were a real (.) tensor)

$\epsilon_{ijk} A^i B^k$ would be a
real vector

Note $\tilde{\epsilon}$ has the same number of
indices as the dimension of space.

$\Rightarrow 4$ for spacetime

\Rightarrow no \times product in 4d "

(but can generalize by contracting
 $\tilde{\epsilon}$ with three vectors A^u, B^v, C^w)